

AB INITIO CALCULATIONS OF LIGHT NUCLEI USING OPTIMIZED CHIRAL HAMILTONIANS

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FUSTIPEN topical meeting: “Understanding Nuclear Structure and Reactions Microscopically, including the Continuum”, Caen, March 17-21, 2014

Main research funding by:



Outline

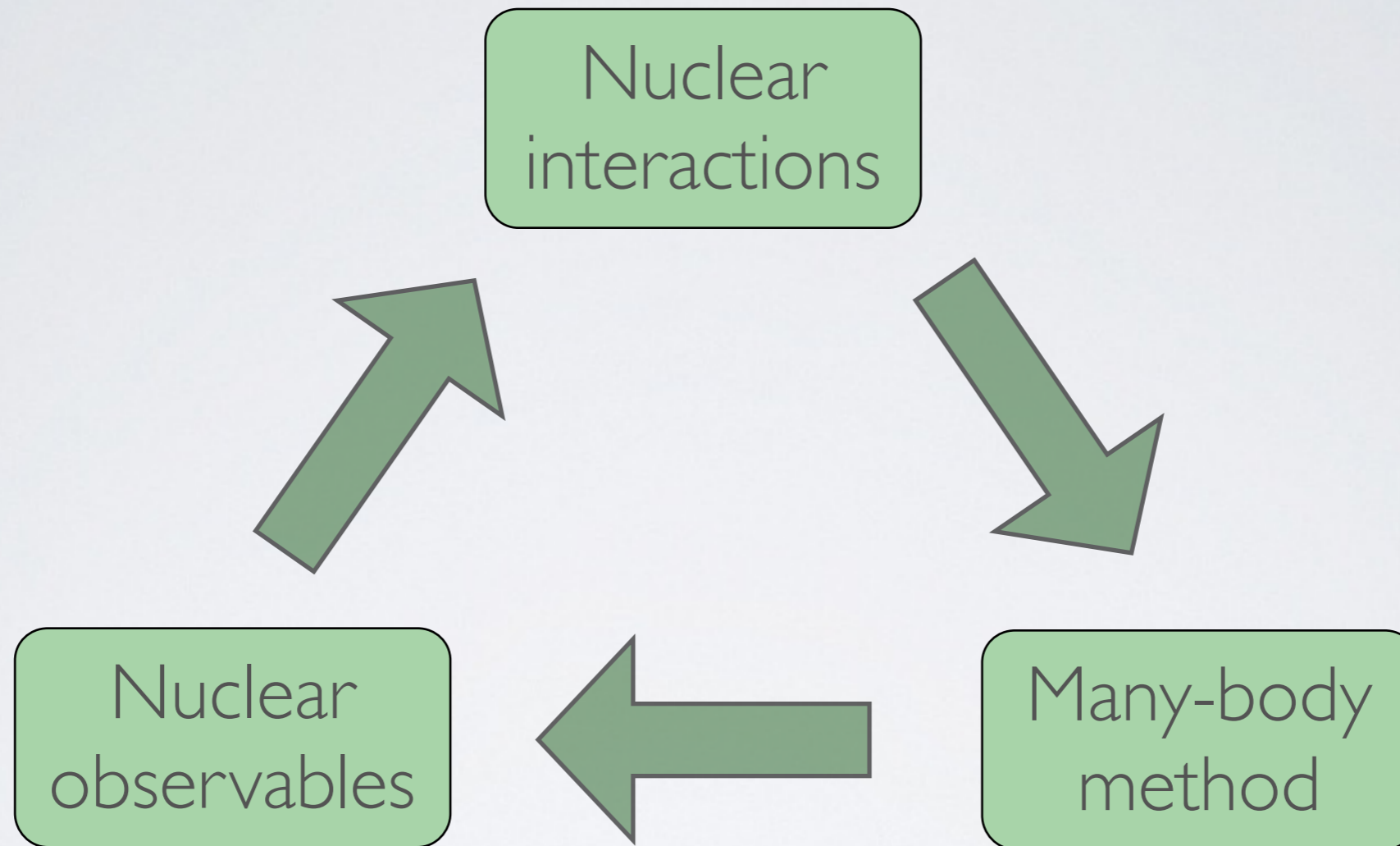
❖ INTRODUCTION

❖ PART I: Optimized chiral Hamiltonians

❖ PART II: From few to many with uncertainty quantification

❖ CONCLUSION

The ab initio nuclear theory approach



Two sources of uncertainty that can be tracked using *ab initio* methods:

- ▶ Nuclear interaction uncertainties
- ▶ Many-body calculation uncertainties

OPTIMIZED CHIRAL HAMILTONIANS



Nuclear forces from chiral EFT

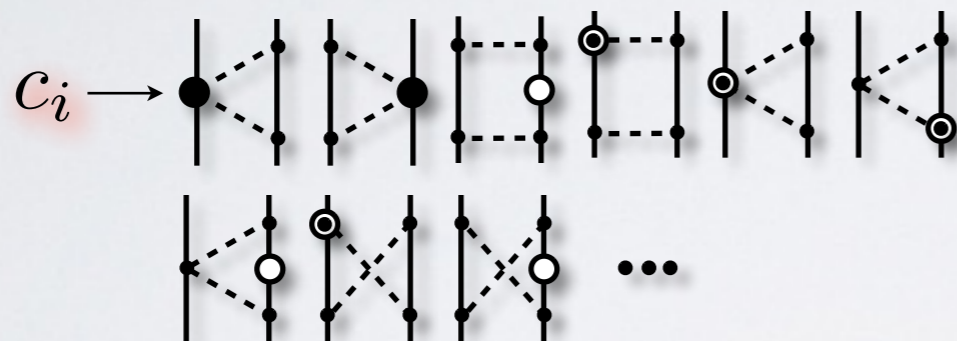
LO



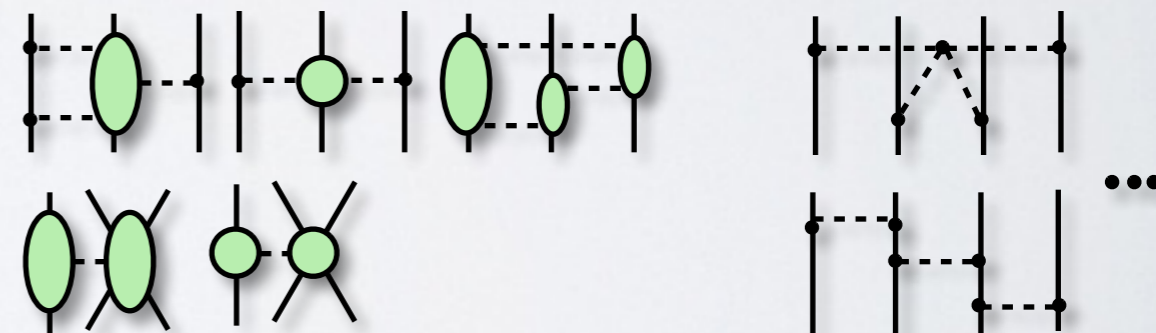
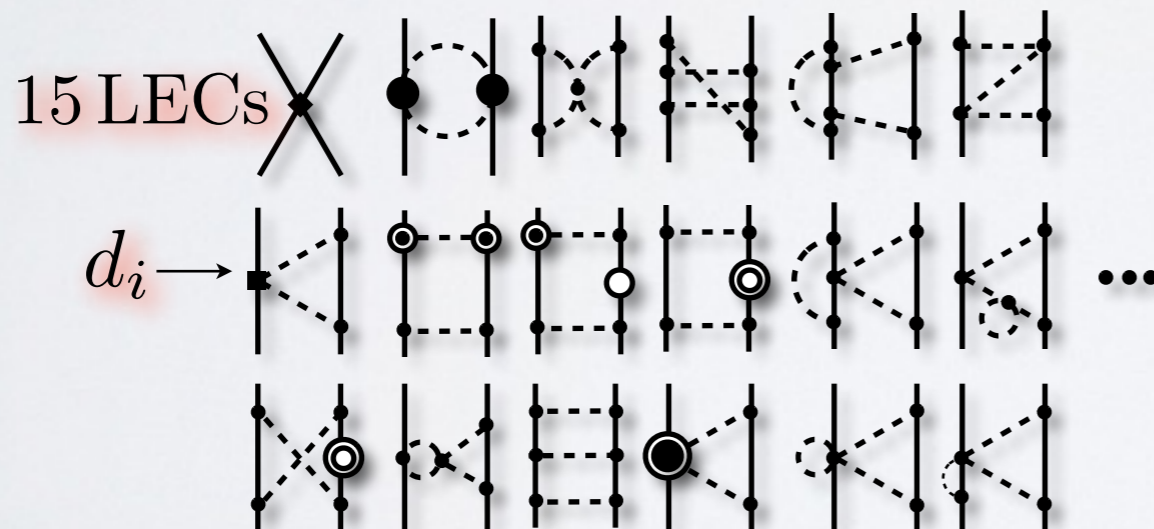
NLO



NNLO



N3LO



⋮

E. Epelbaum et al. Rev. Mod. Phys. 81, 1773 (2009)
R. Machleidt et al. Phys. Rep. 503, 1 (2011)

C. Forssén, Caen, Mar. 21, 2014



Low-energy constants at NNLO

Optimized chiral NN interaction at NNLO

• A. Ekström, et al., 2013. Phys. Rev. Lett., **110** (2013), 192502.



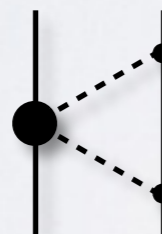
2 LECs

$$\tilde{C}_{1S_0}^{pp} \quad \tilde{C}_{1S_0}^{np} \quad \tilde{C}_{1S_0}^{nn} \quad \tilde{C}_{3S_1}$$



7 LECs

$$C_{1S_0} \quad C_{3P_0} \quad C_{3P_1} \quad C_{3P_2} \\ C_{1P_1} \quad C_{3S_1} \quad C_{3S_1} - {}^3D_1$$



$\dots C_i$

$$C_1 \quad C_3 \quad C_4$$

long-range physics



C_i

C_D

C_E

two new parameters that can be determined from the $A > 2$ systems



Two new optimization strategies

Any “potential model” will contain coupling constants that need to be determined from data one way or the other.

$$\min_{\vec{x}} \left[f(\vec{x}) = \sum_{q=1}^N \left(\frac{O(\vec{x})_q - O_q^{\text{exp}}}{w_q} \right)^2 \right]$$

Define an objective function that is relevant for the “objective”

1. **π N LECs determined first** from Pion-Nucleon scattering phase shifts
2. **(NN-only) objective function based on observables only**
 - ▶ Nucleon-Nucleon scattering data
 - ▶ 1S_0 effective range expansion
 - ▶ A=2: energy, radius, Q
3. **NNN LECs determined at the end** given the NN part

1. **Expanded (NN+NNN) objective function**
 - ▶ Nucleon-Nucleon scattering data
 - ▶ Pion-Nucleon scattering phase shifts
 - ▶ 1S_0 effective range expansion
 - ▶ Few-body observables: NCSM 2H/3H/4He Binding energies and radii
 - ▶ Energy and saturation momentum of symmetric nuclear matter from MBPT2



πN LECs: overview

	piN-Krebs	piN-BM	NN-PWA	NNLO (Bochum)	N3LO (Idaho)	NNLOopt
c1	[-1.13,-0.75]	-0.81±0.12	-0.76±0.07	-0.81	-0.81	-0.9186
c3	[-5.51,-4.77]	-4.70±1.16	-4.78±0.10	-3.4	-3.2	-3.8887
c4	[3.34,3.71]	3.40±0.04	+3.96±0.22	+3.40	+5.40	+4.3103

piN-Krebs:

The most recently published, and to fourth order, analysis of the piN scattering phase shifts up to pLab=150 MeV [GW06,KH86]

piN-BM:

Analysis of KA84 piN scattering phase shifts pLab=40-97

NN-PWA:

Nijmegen PWA analysis of NN scattering data, with the long range physics described by subleading chiral two-pion exchanges

NNLO (Juelich):

pion-nucleon couplings taken from piN-BM, but c3 chosen on the larger side within the uncertainty. This value is consistent with the Entem Machleidt analysis of NN data.

N3LO (Idaho):

Guided by fit to NN data

NNLOopt:

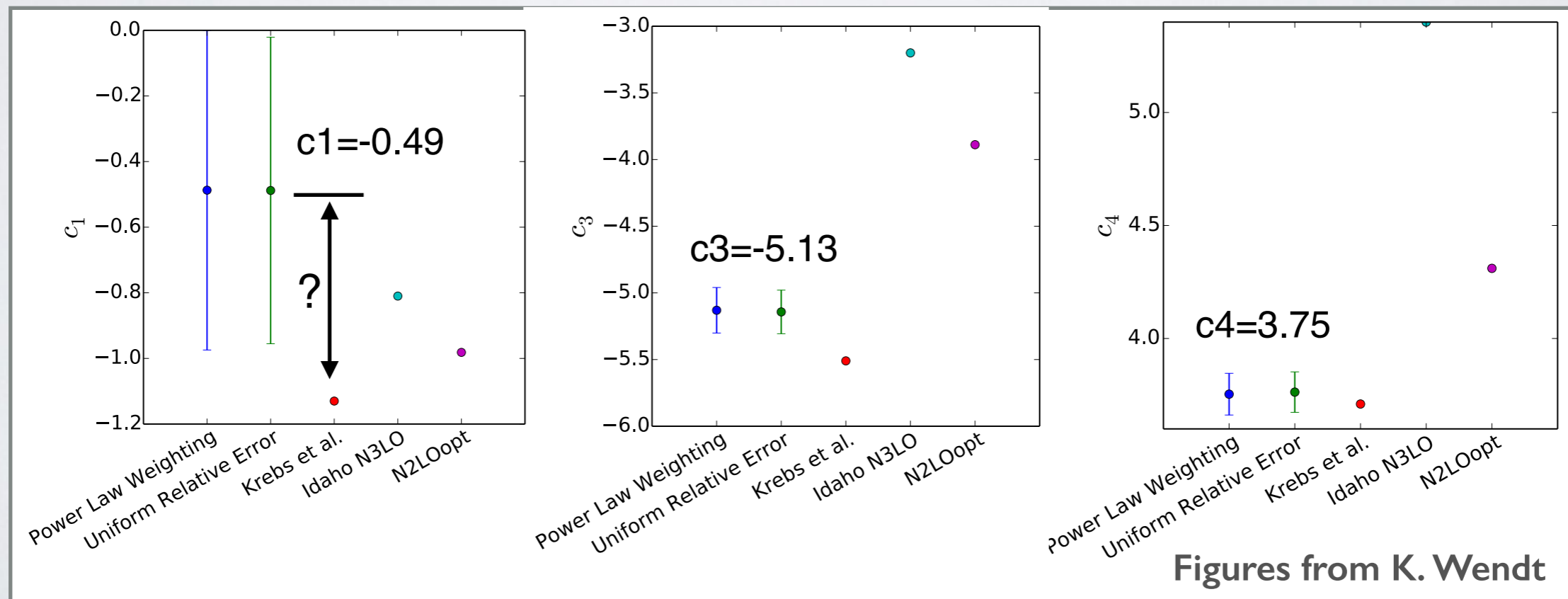
Guided by fit to NN data

- E. Epelbaum *et al.* Rev. Mod. Phys. 81, 1773 (2009)
- E. Epelbaum *et al.* Eur. Phys. J. A19, 401 (2004)
- R. Machleidt *et al.* Phys. Rep. 503, 1 (2011)
- A. Ekström *et al.* Phys. Rev. Lett. 110, 192502 (2013)
- H. Krebs *et al.* Phys. Rev. C, 85, 054006 (2012)
- M. C. M. Rentmeester *et al.* Phys. Rev. C, 67, 044001 (2003)



πN LECs: new results

	piN-Krebs	piN-BM	NN-PWA	NNLO (Bochum)	N3LO (Idaho)	NNLOopt
c1	[-1.13,-0.75]	-0.81±0.12	-0.76±0.07	-0.81	-0.81	-0.9186
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order-by-order optimization

In progress

❖ LO optimization

- ▶ NN scattering observables ($T_{\text{lab}} \leq 0.5$ MeV), 1S_0 ERE, $A=2$ observables

❖ NLO optimization

- ▶ NN scattering observables ($T_{\text{lab}} \leq 75$ MeV, s-waves / p-waves mainly for 0-8 / 8-75 MeV, respectively), 1S_0 ERE, $A=2$ observables

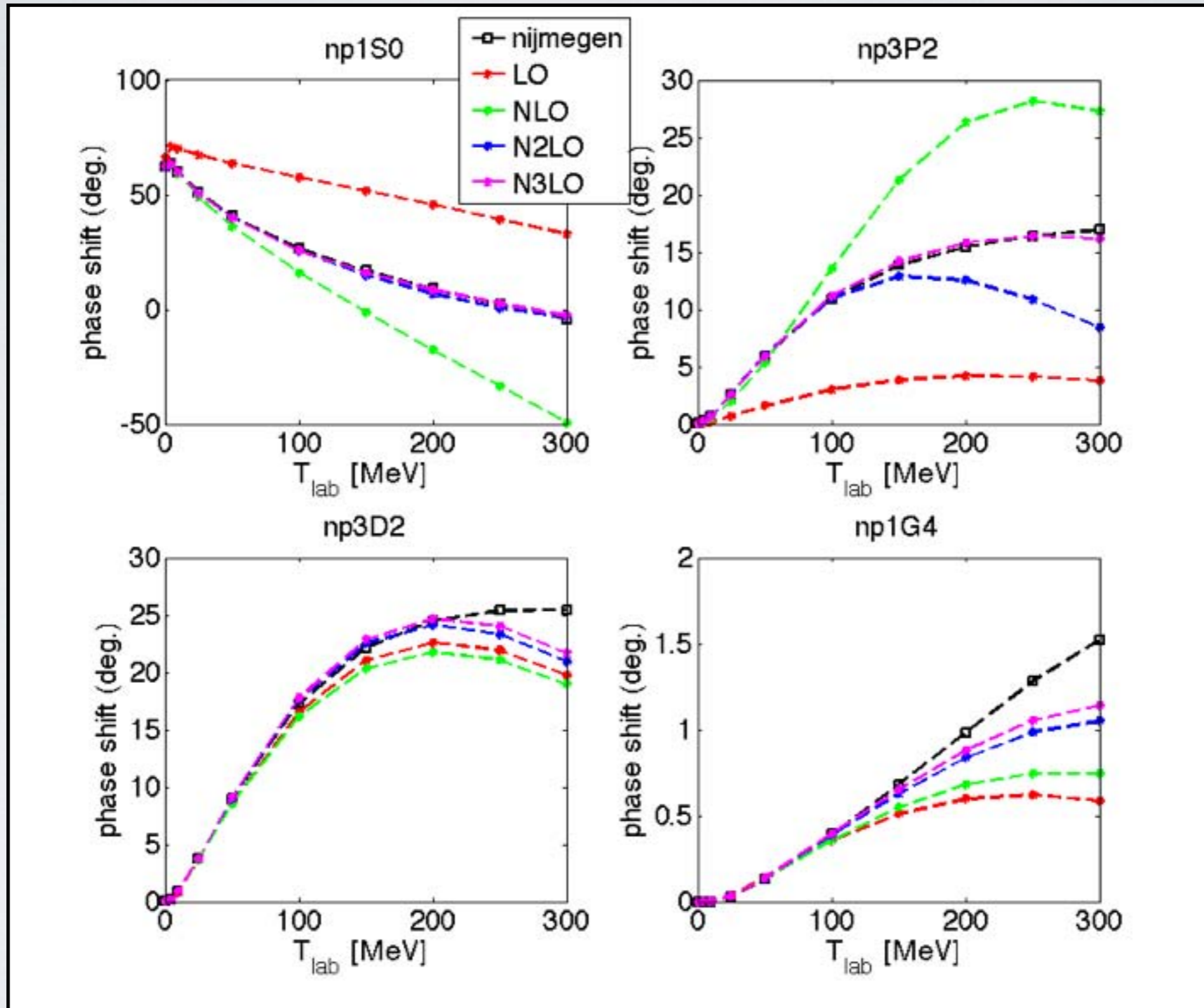
❖ NNLO optimization

- ▶ Work in progress. The results presented here are with NNLO_{opt}

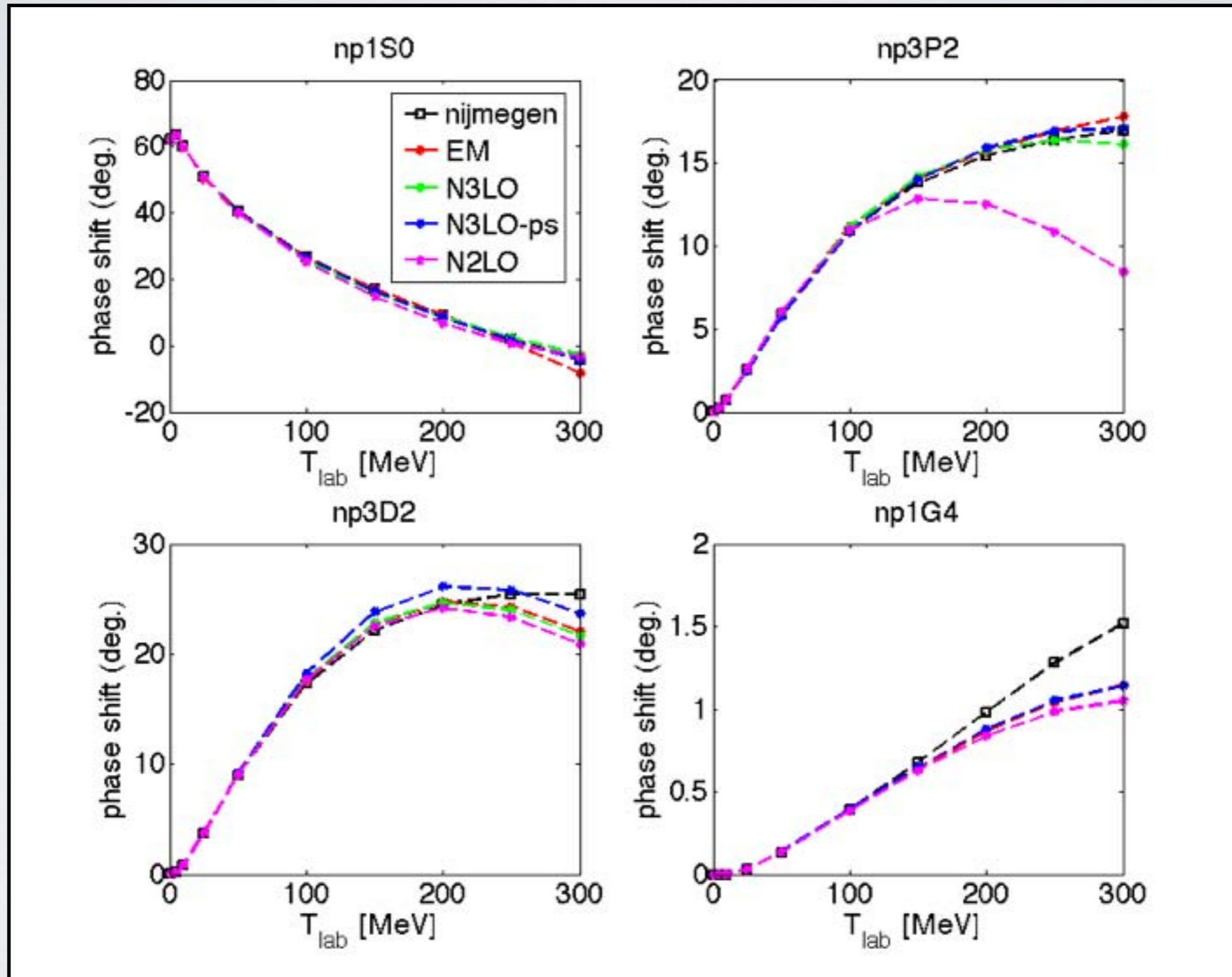
❖ N3LO optimization

- ▶ NN scattering observables ($T_{\text{lab}} \leq 350$ MeV), 1S_0 ERE, $A=2$ observables

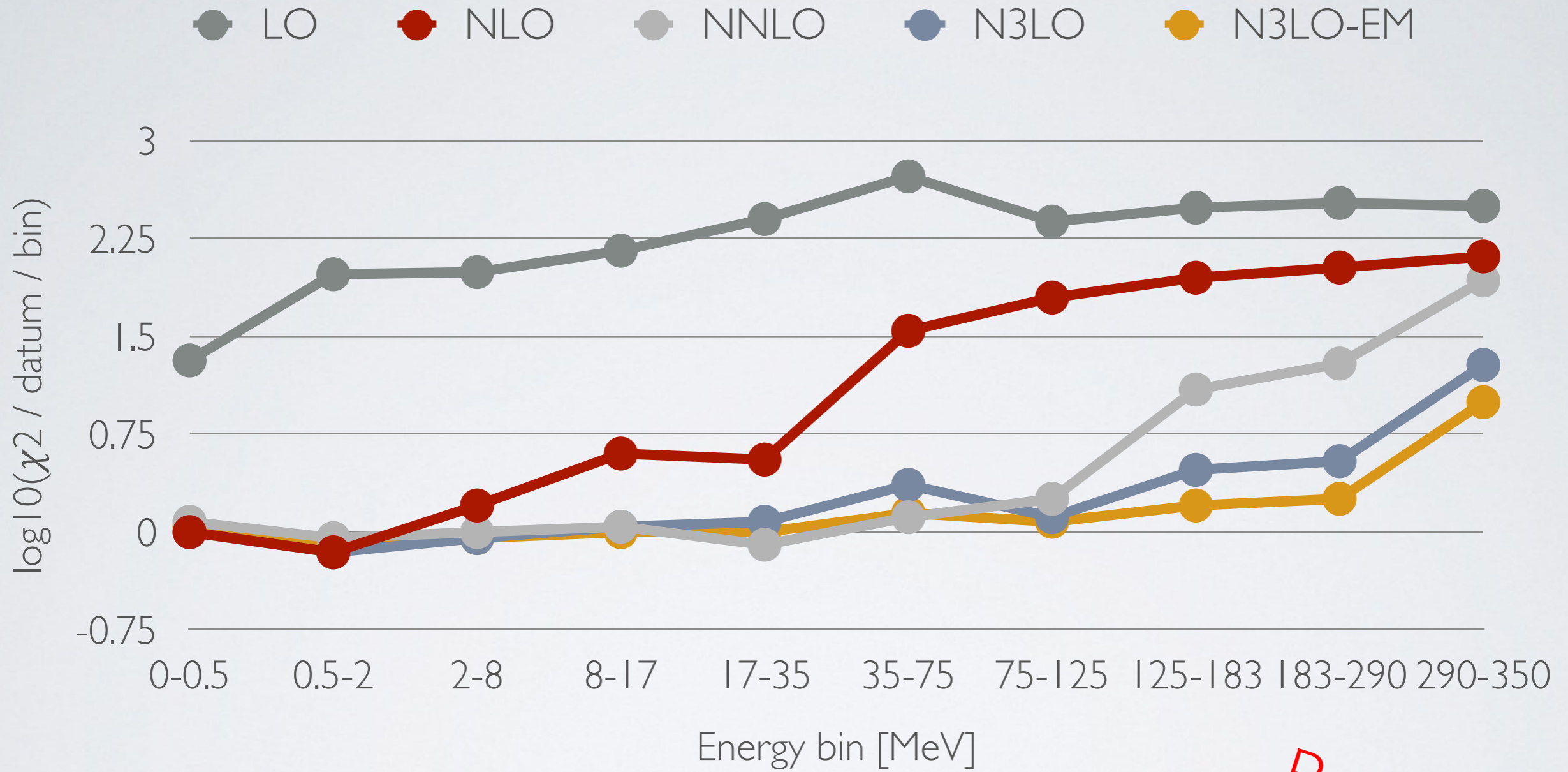
Selected np phase shifts



Selected np phase shifts



χ^2 / datum per energy bin



Preliminary



FROM FEW TO MANY



Frontiers of *ab initio* nuclear many-body theory

- ▶ The **many-body frontier**
 - ▶ Heavier systems, away from closed shells
- ▶ The **continuum frontier**
 - ▶ Approaching the drip lines
 - ▶ Unified theory of structure and reactions
- ▶ The **convergence frontier**
 - ▶ Uncertainty quantification, error propagation



The No-Core Shell Model

❖ Many-body Schrödinger equation

▶ A-nucleon wave function

▶ Non-relativistic, point nucleons

❖ Hamiltonian:

$$H_A = \frac{1}{A} \sum_{i < j}^A \frac{(\vec{p}_i - \vec{p}_j)^2}{2m} + \sum_{i < j}^A V_{NN,ij} + \sum_{i < j < k}^A V_{NNN,ijk}$$

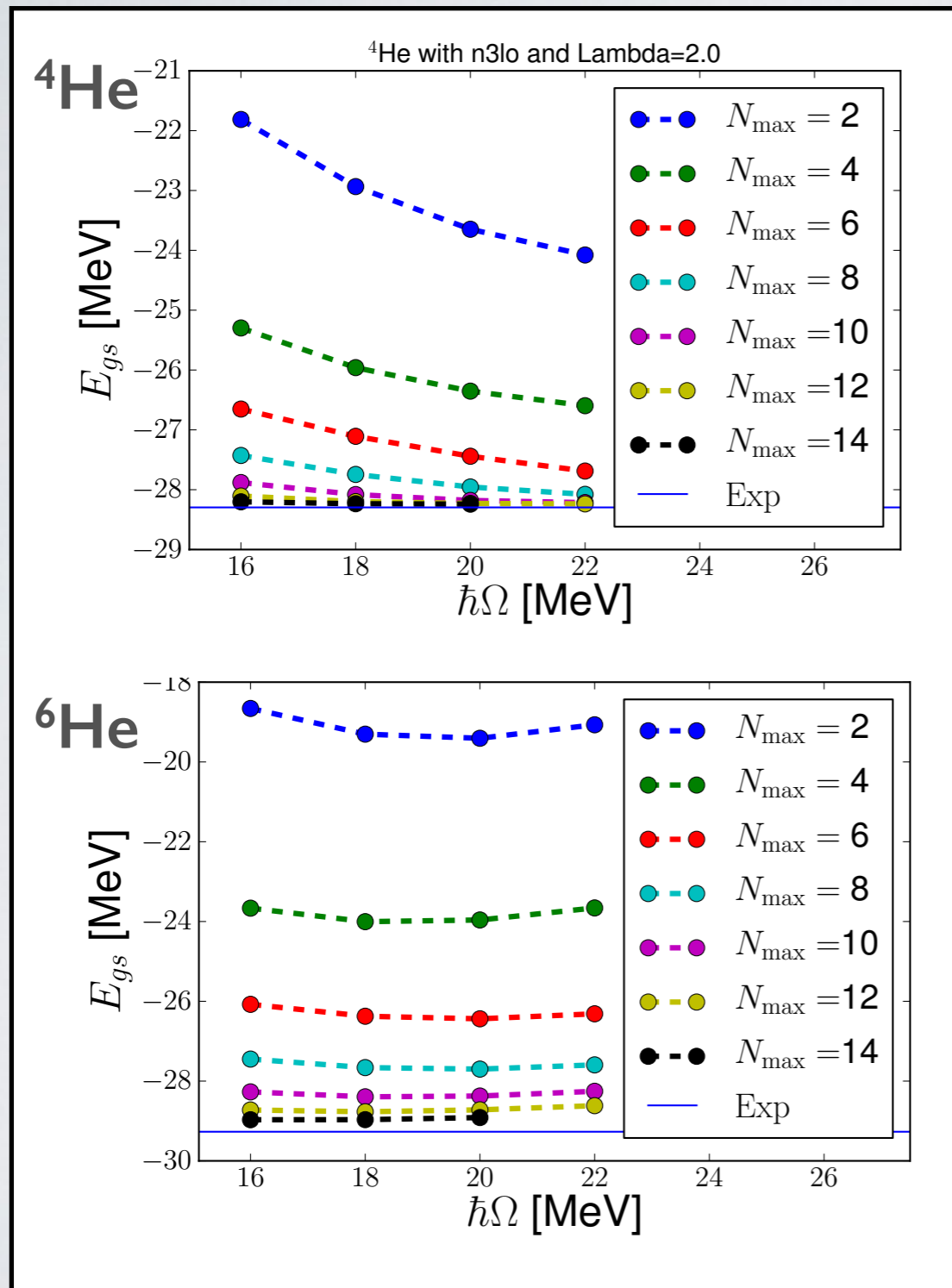
❖ Many-body basis: Slater determinants composed of harmonic oscillator single-particle states

❖ Respects translational invariance and includes full antisymmetrization



NCSM convergence: finite HO space

$N^3\text{LO}$, SRG (NN only, $\Lambda = 2.0 \text{ fm}^{-1}$)



HO basis cutoff scales

$$\Lambda_{\text{UV}} = \sqrt{2(N + 3/2)}\hbar/b$$

$$L_{\text{IR}} = L_2 \equiv \sqrt{2(N + 3/2 + 2)}b$$

Extrapolations from finite HO basis

- R.J. Furnstahl et al., Phys. Rev. C 86(2012)031301R
- S. Coon et al., Phys. Rev. C 86(2012)054002
- R.J. Furnstahl et al., Phys. Rev. C 87(2013)044326
- R.J. Furnstahl et al., arXiv:1312.6876

Previous work with $N_{\text{max}} / \hbar\Omega$ extrapolation

- C. Forssén et al., Phys. Rev. C 77(2008)024301
- P. Maris et al., Phys. Rev. C 79 (2009)014308

LO correction to the energy due to finite HO space (Dirichlet bc):

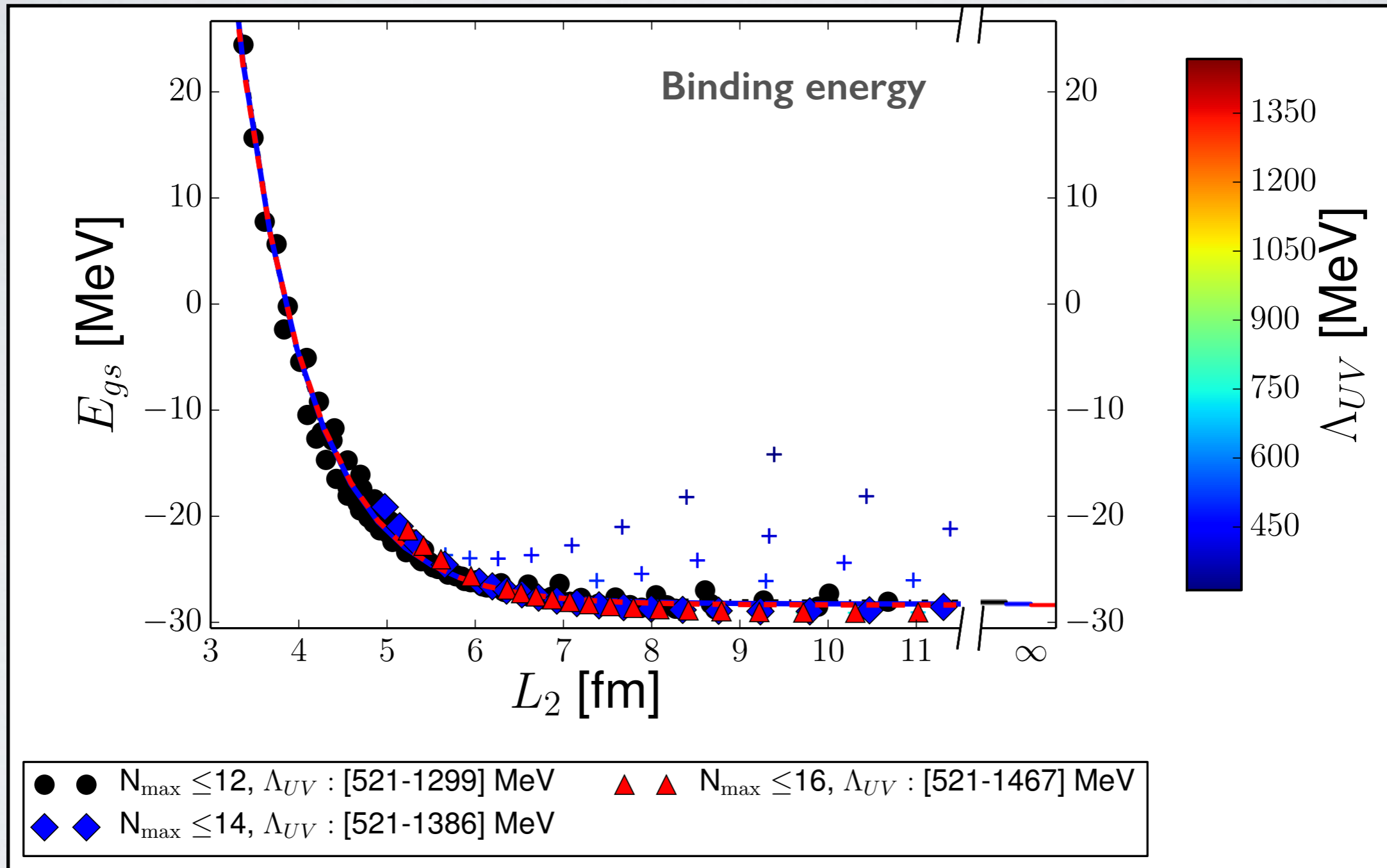
$$E_L = E_\infty + \Delta E_L$$

$$\text{with } \Delta E_L = a_0 \exp(-2k_\infty L)$$

HO basis extrapolation schemes

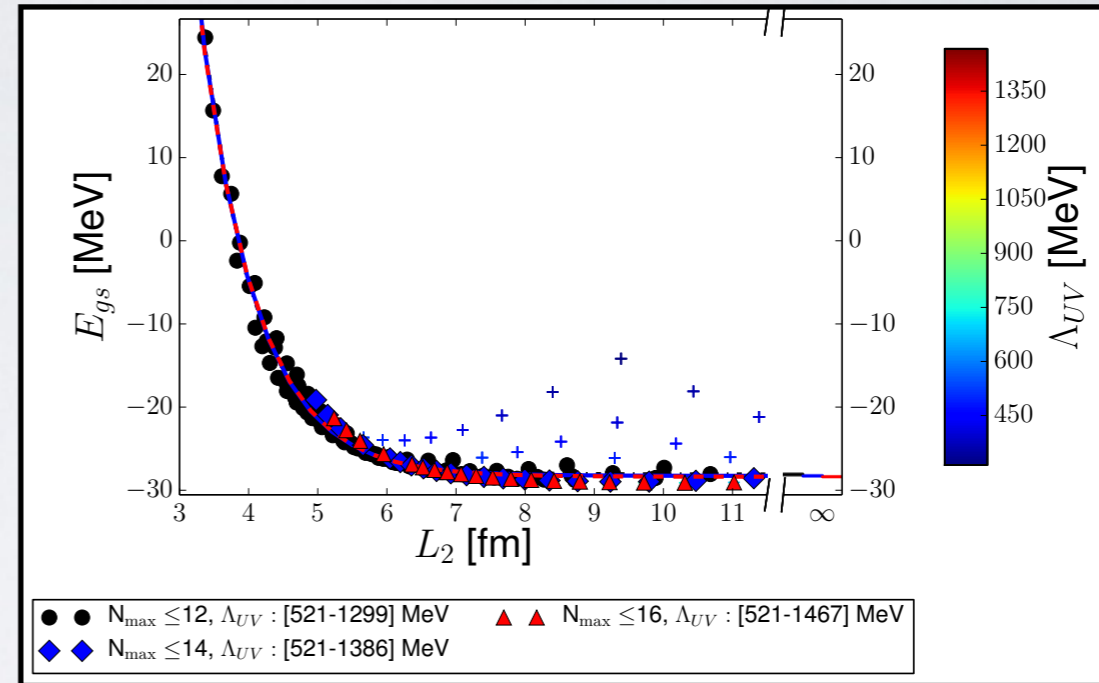
${}^6\text{He}$

N^3LO , SRG (NN only, $\Lambda = 2.0 \text{ fm}^{-1}$)



${}^6\text{He}$: Ground-State Properties

$N^3\text{LO}$, SRG (NN only, $\Lambda = 2.0 \text{ fm}^{-1}$)



	Exp. ⁵	This work			Bacca <i>et al.</i> ⁶
		$\Lambda_{\text{SRG}} = 1.8$	$\Lambda_{\text{SRG}} = 2.0$	$\Lambda_{\text{SRG}} = 2.2$	$V_{\text{low-}k} (\Lambda = 2.0 \text{ fm}^{-1})$
E_{gs} [MeV]	29.269	29.42(9)	28.79(6)	28.03(10)	29.47(3)
S_{2n} [MeV]	0.975	0.97(9)	0.54(6)	0.09(11)	0.82(4)
$r_{\text{pt-p}}$ [fm]	1.938(23)	1.786(4)	1.788(3)	1.805(4)	1.804(9)



CONCLUSION



Summary

- ❖ Optimization of chiral Hamiltonian up to N³LO
 - ▶ POUNDerS shows great promise to deliver state-of-the-art optimized nuclear forces
 - ▶ Optimized with respect to scattering data
 - ▶ Two different optimization strategies
- ❖ No-core shell model for light nuclei
 - ▶ Introduction of UV and IR scales gives error estimate
- ❖ Outlook
 - ▶ Correlated uncertainty estimates
 - ▶ Error propagation to nuclear observables
 - ▶ Cutoff dependence



Acknowledgements

Many thanks to my collaborators

- ❖ Jimmy Rotureau, Emil Ryberg, **Daniel Sääf**, **Boris Carlsson**, Håkan Johansson (Chalmers), Lucas Platter (ANL), **Andreas Ekström** (UiO, Chalmers)
- ❖ Petr Navrátil (TRIUMF), Robert Roth (Darmstadt), James Vary and Pieter Maris (Iowa), and the NCSM collaboration.
- ❖ Marek Ploszajczak and Nicolas Michel (GANIL), George Papadimitriou (Iowa), and Bruce Barrett (UA)
- ❖ Morten Hjorth Jensen (UiO, MSU), Witek Nazarewicz, Gaute Hagen, Gustav Jansen, Thomas Papenbrock, **Kyle Wendt** (ORNL/UT), Ruprecht Machleidt (Idaho), Jason Sarich and Stefan Wild (ANL)
- ❖ Nikolai Zinner, Artem Volosniev (Aarhus)



Research funded by:

- The Swedish Research Council
- European Research Council
- STINT

